# Current problems in cosmic ray physics

Taavi Tuvi

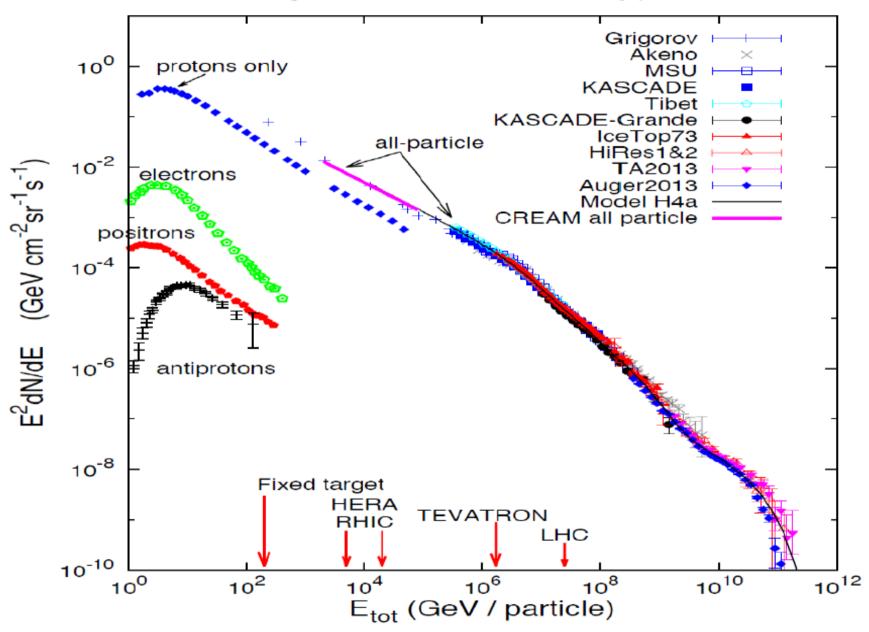
### Contents

Cosmic rays "standard model"

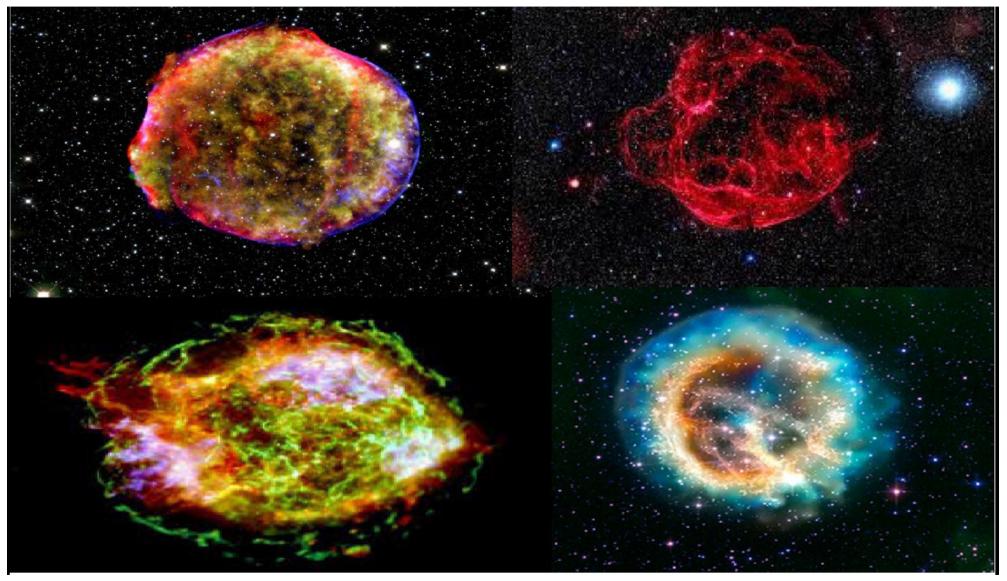
What the current data shows

## Cosmic rays

- Cosmic rays are high energy charged particles, arriving to Earth from interstellar space
- Their energy spans from a few hundred MeV up to 300EeV (3x10<sup>20</sup> eV).
- Main component is nuclei, almost 99%
  - Of nuclei about 89% are protons
  - About 10% helium nuclei
  - 1% heavier nuclei
- 1% of rest is mainly electrons

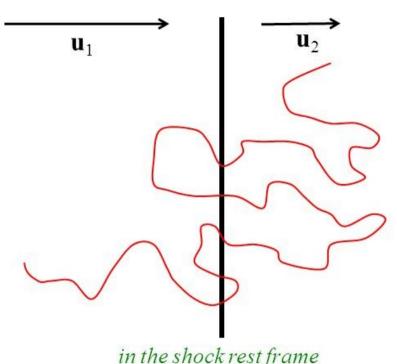


## Production



10.5.17

## Diffusive shock acceleration: $r_g >> d$



shock compression

$$R \oplus u_1/u_2$$

I order acceleration

$$\Delta p \sim p \, \frac{u}{v}$$

where u = u1-u2

Compressive discontinuity of the plasma flow leads to acceleration of particles reflecting at both sides of the discontinuity: diffusive shock acceleration (I-st order Fermi)

the differential energy spectrum is:

compression ratio 
$$r = u_1/u_2$$

$$n \mathrm{d} p \propto p^{-(r+2)/(r-1)} \mathrm{d} p.$$

## Propagation

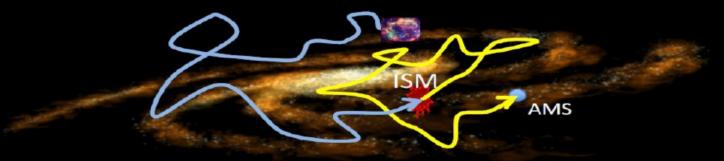
#### Primary Cosmic Rays (p, He, C, O, ...)



Primary cosmic rays carry information about their original spectra and propagation.

#### Secondary Cosmic Rays (Li, Be, B, ...)

C, O, ..., Fe + ISM  $\rightarrow$  Li, Be, B + X



Secondary cosmic rays carry information about propagation of primaries, secondaries and the ISM.

10

## Propagation

Galactic magnetic field's turbulence spectrum

- Kolmgorovs delta 1/3
- Kraichnan delta 1/2

 Secondaries – grammage of material traversed, function of CR residence time in the galaxy

## Precision CR physics

Pamela (2008) – discovery of the positron excess

 AMS-02 (2011) – measurment of CR fluxes up to 1%

## 5 years of AMS



## CERN Colloquium

SPEAKER: Prof. TING, S. (Massachusetts Inst. of

Technology (US))

TITLE: The First Five Years of the Alpha

**Magnetic Spectrometer on the** 

**International Space Station** 

DATE: Thu 08/12/2016 17:00

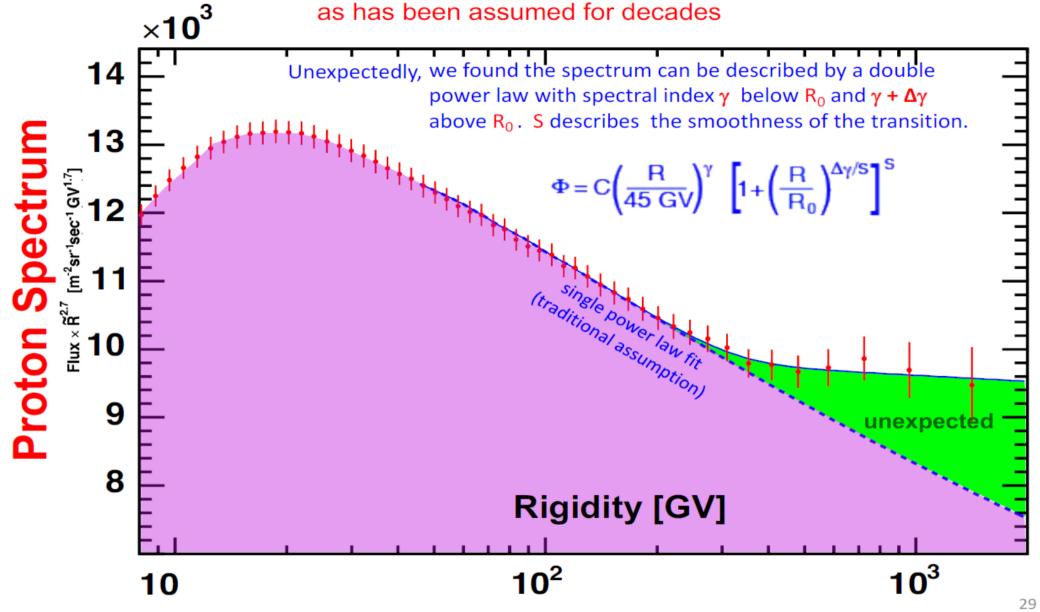
# Unexpected Excesses in everything

CR standard model predicts

- Power law sources
- Power law preserving transport
- Power law fluxes at Earth

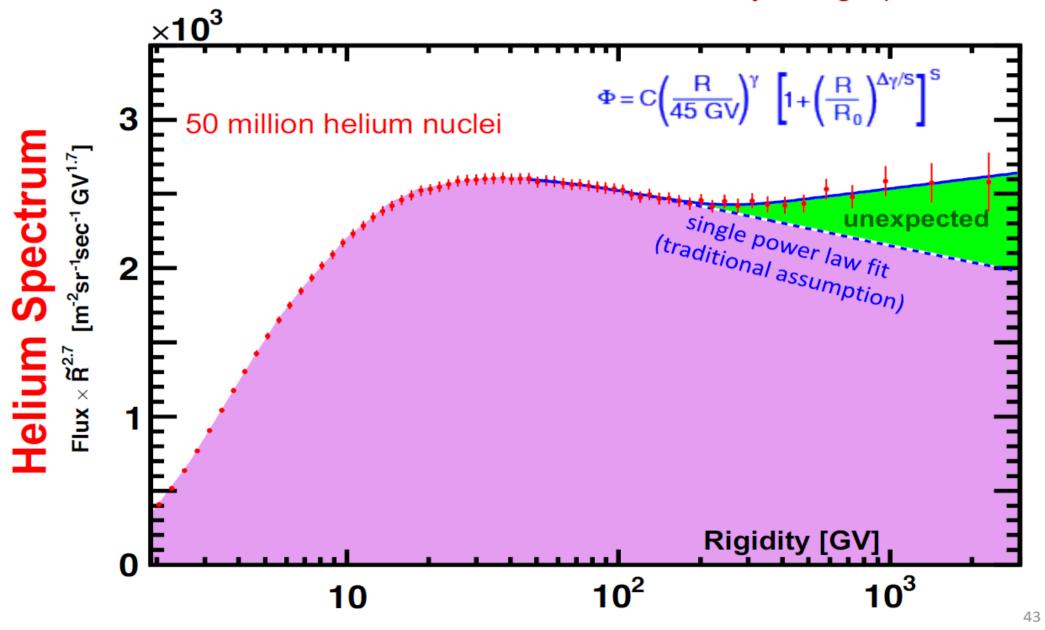
#### **AMS** proton flux

New information: The proton flux cannot be described by a single power law =  $CR^{\gamma}$ , as has been assumed for decades



#### **AMS Helium Flux**

New information: The Helium flux cannot be described by a single power law.

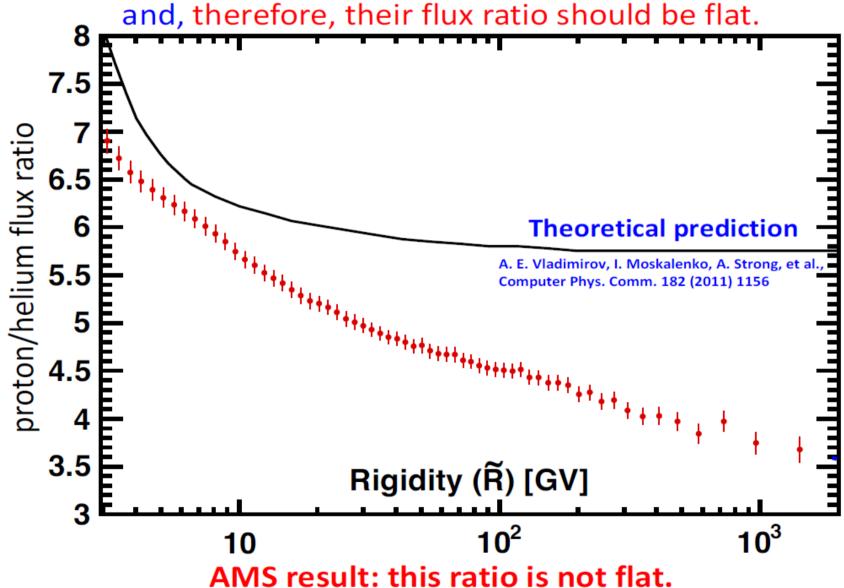


10.5.17 Slide by Prof. Samuel Tiħogavere
The First Five Years of the Alpha Magnetic Spectrometer on the International Space Station
CERN Colloquium 8 Dec 2016

#### Physics Result 9: The AMS proton/helium flux ratio

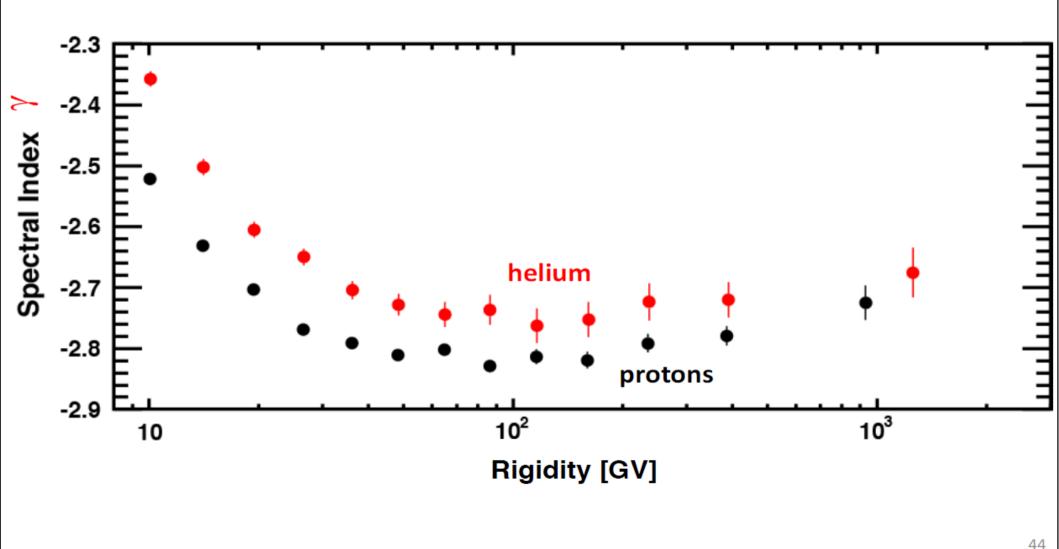
Protons and helium are both "primary" cosmic rays.

Traditionally, they are assumed to be produced in the same sources



45

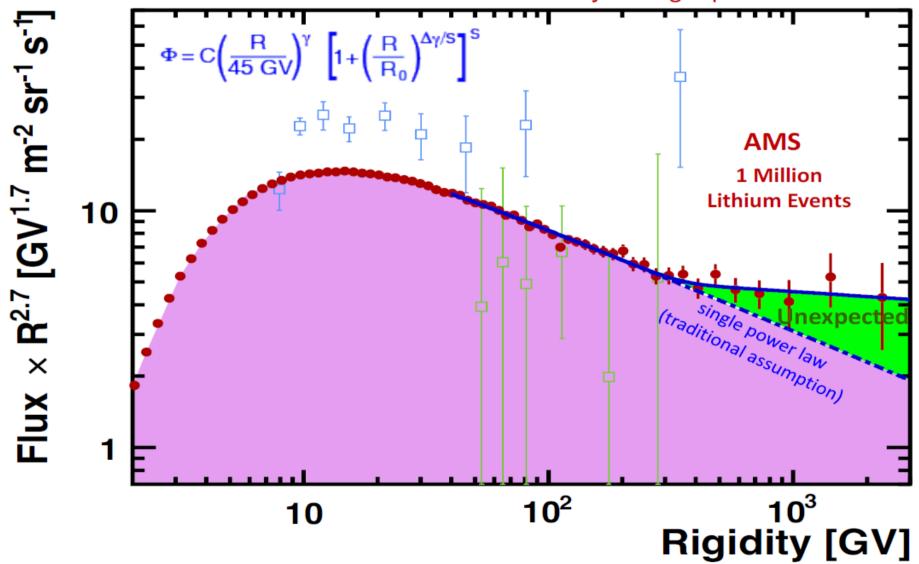
## New information: The helium spectral index changes with rigidity in a similar way to that of the proton spectral index but the values are different



#### Physics Result 10: The Lithium flux

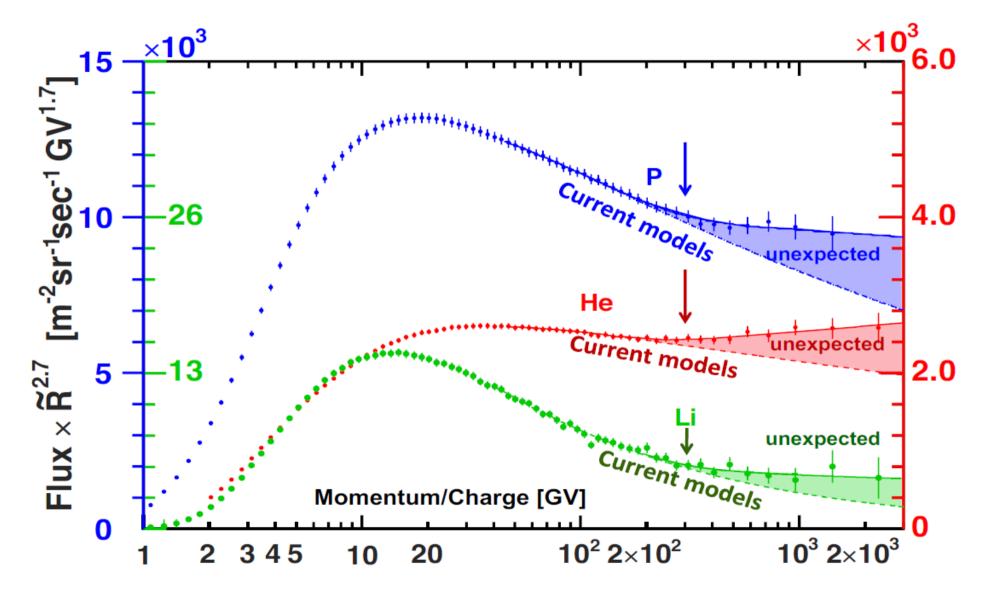
#### New AMS results on Secondary Cosmic Rays (Lithium)

New information: The Lithium spectrum behaves similar to protons and Helium and the Lithium flux cannot be described by a single power law.



#### **Summary (on nuclei)**

The spectra of protons, helium and lithium do not follow the traditional single power law. They all change their behavior at the same energy.



17

CR nuclei have a break at rigidity ~450GV

Proton and helium fluxes have different indexes

Propagation?

## Low energy break

- CR nuclei seem to have a break at rigidity ~10-20GV
- With low energy spectral index being ~2.3-2.4
- That being consistent with diffusive gamma-ray background from inner galaxy

"Cosmic ray spectrum in the local Galaxy"

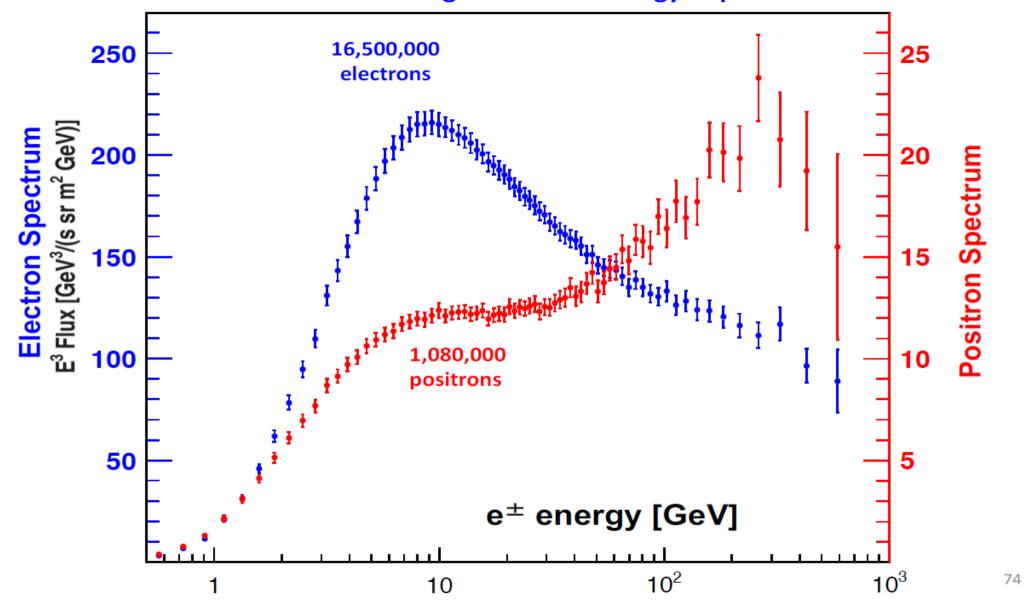
Andrii Neronov, Denys Malyshev, Dmitri V. Semikoz

https://arxiv.org/abs/1705.02200

### Antimater excesses

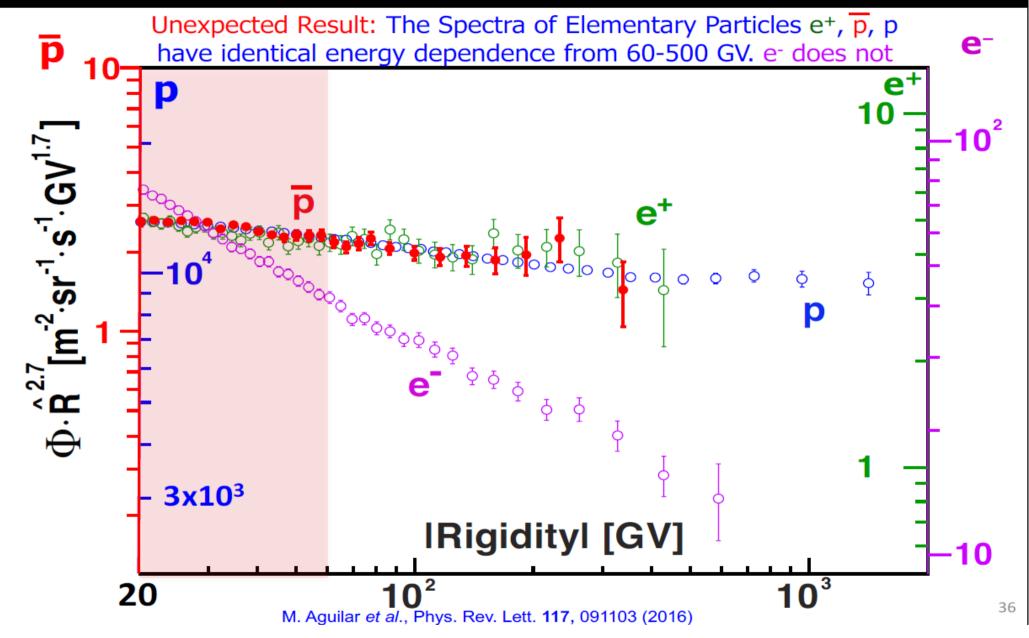
#### **Summary (on elementary particles)**

The electron flux and the positron flux are different in their magnitude and energy dependence.



10.5.17 Slide by Prof. Samuel Tiħogavere
The First Five Years of the Alpha Magnetic Spectrometer on the International Space Station
CERN Colloquium 8 Dec 2016

## Physics Result 7: The antiproton flux and properties of elementary particle fluxes



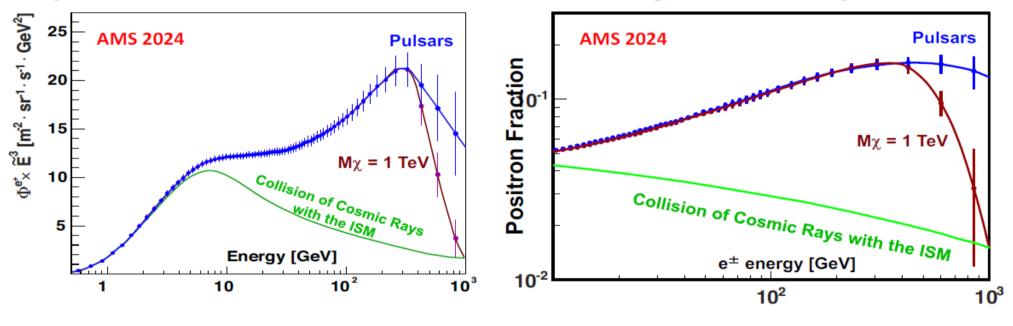
10.5.17 Slide by Prof. Samuel Tiħogavere
The First Five Years of the Alpha Magnetic Spectrometer on the International Space Station
CERN Colloquium 8 Dec 2016

#### **Summary (on elementary particles)**

#### **Positron Spectrum**

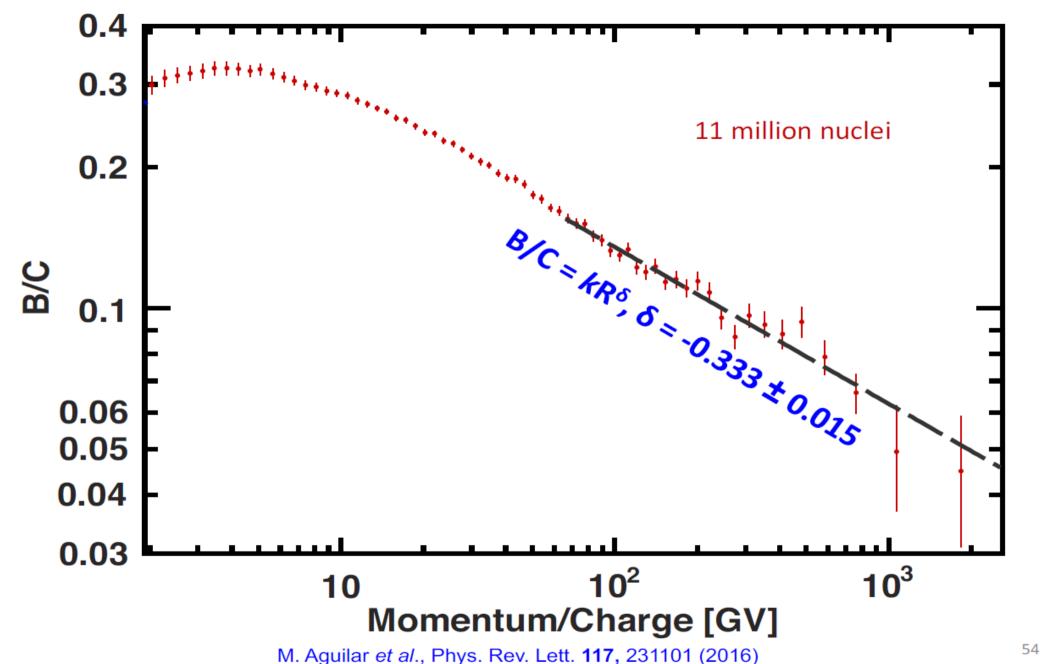
#### **Positron Fraction**

By 2024 we will should be able understand the origin of this unexpected data.



## Some clarity with propagation

#### Physics Result 14: The Boron-to-Carbon (B/C) flux ratio



10.5.17 Slide by Prof. Samuel Tiñgavere The First Five Years of the Alpha Magnetic Spectrometer on the International Space Station CERN Colloquium 8 Dec 2016

#### **AMS B/C results**

The B/C ratio does not show any significant structures in contrast to many cosmic ray models that require such structures at high rigidities.

Remarkably, above 65 GV, the B/C ratio is well described by a single power law B/C =  $k R^{\delta}$  with  $\delta = -0.333 \pm 0.015$ .

> This is in agreement with the Kolmogorov turbulence model of magnetized plasma of  $\delta = -1/3$  asymptotically.

## Summary

Possible explenations

Dark matter

Flawed models of producton and propagation

Local effects

## Thank you for your attention!